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A Bayesian Hierarchical Approach to Ensemble Weather Forecasting

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Abstract

In meteorology, the traditional approach to forecasting employs deterministic models mimicking atmospheric dynamics. Forecast uncertainty due to the partial knowledge of initial conditions is tackled by Ensemble Predictions Systems (EPS). Probabilistic forecasting is a relatively new approach which may properly account for all sources of uncertainty. In this work we propose a hierarchical Bayesian model which develops this idea and makes it possible to deal with an EPS with non-identifiable members using a suitable definition of the second level of the model. An application to Italian small-scale temperature data is shown.

Keywords. Ensemble Prediction System, hierarchical Bayesian model, predictive distribution, probabilistic forecast, verification rank histogram.

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1 Introduction

In meteorology, the typical approach to forecasting employs deterministic models mimicking atmospheric dynamics. Probabilistic forecasting is a relatively new approach attracting growing interest (see Gneiting et al., 2007). One technique currently used to tackle the problem of weather predictability is Ensemble Weather Forecasting. A prediction system based on this technique is called an Ensemble Prediction System (EPS). An EPS produces multiple weather forecasts by iterating forward random perturbations of a best estimate of initial conditions (Leith, 1974; Toth et al., 2001).

Weather forecasts obtained using an EPS can be synthesised to give one single value. Such a value may derive from an average, as in Raftery et al. 2005, or from a selection procedure, as in Roulston and Smith 2003, where the Best Member Dressing method is proposed. One important distinguishing feature of an EPS is the identifiability of its outputs. The competing models of Raftery et al. 2005 are indeed different forecasting models which may, however, agree upon a common evaluation. The model used in Roulston and Smith 2003 is completely different: in fact, this uses the EPS from the European Center for Medium-Range Weather Forecasts (ECMWF; Montani et al., 1996), where model outputs cannot be identified since they are replicas under perturbation of initial conditions. A number of different ways of treating EPS meteorological forecasts have been suggested.

The Bayesian Model Averaging (BMA; as in Hoeting et al., 1999) approach adopted by Raftery et al. 2005 focuses on one application to the University of Washington’s multi-model EPS. It treats forecasts from deterministic models as inputs for a statistical model and can be extended to dynamical models. Of the forecasts $f_k$, each computed according to model $k$, there is a ‘best’ forecast, while the $K$ models are
different and identifiable. Uncertainty about which model is the best one is naturally quantified by BMA. Denote as $y$ the observed value, and each deterministic forecast as $\hat{f}_k$. Each deterministic forecast can be corrected for possible bias, for example by a linear transformation, thus giving the corrected forecast $f_k = a_k + b_k \hat{f}_k$. A conditional predictive distribution function $g_k(y|f_k, \theta)$ is associated to $f_k$ and this function can be interpreted as the distribution of $y$ conditional on $f_k$, given that $f_k$ is the best forecast in the ensemble. The predictive distribution function as given by BMA is (cfr. Raftery et al., 2005, eq. (2)):

$$p(y|f_1, f_2, \ldots, f_K) = \sum_{k=1}^{K} w_k g_k(y|f_k, \theta)$$  \hspace{1cm} (1)

where $w_k$ is the probability that $f_k$ is the best forecast, and is based on the predictive performance of the $k$-th dynamic model in a training set of deterministic forecasts and observed values. Weights $w_k$ are model probabilities, and sum to 1. Estimation of parameter $\theta$ is performed by ML, with the EM algorithm. Expression (1) is referred to as Bayesian since it weighs likelihoods by evaluating the appropriateness of each model, which in this context can be seen as a ‘state of the world’.

Roulston and Smith 2003 propose the Best Member Dressing method, and analyse the ECMWF EPS. Given an ensemble forecast, it is unlikely that any of the ensemble members will equal the observation. The lack of correspondence can be accounted for by assigning an error distribution to each ensemble member. In order to do so, one needs to know the appropriate degree of uncertainty. Roulston and Smith’s proposal associates uncertainty with the ensemble’s best member, where the best member is defined as the one nearest to the observed data in the weather system state space. One application to European ECMWF EPS is shown regarding 4 stations: Tromso, London Heathrow, Frankfurt and Warsaw. Statistical ensembles showed a significant edge in performance over deterministic ensembles. However, substantial residual variability remains, and this cannot be explained by EPS.
The present study starts from the premise that observed data and EPS outputs may be perceived as the components of a statistical model in which observed meteorological values are explained by EPS outputs considered as exogenous variables, as in Raftery et al. 2005. Our model completes the idea of the BMA in two ways. Firstly, it is conceived as a genuinely Bayesian model, and secondly, it deals with non-identifiable EPS outputs such as those present in the European ECMWF EPS, by modelling the second level of the hierarchy in an appropriate manner. In meteorology, observed values do not constitute current inputs for meteorological models and the comparison between observed and predicted values is not as crucial as it can be in other contexts (Gneiting et al., 2007). After integrating out all parameters, our model can be used to obtain the distribution of an as yet unobserved value, conditional on previous observations and forecasts. Our model is not a tool for improving forecasts, but associates probability distributions to EPS outputs, and can explain the residual variability which is not taken into account by an EPS. It obtains truly probabilistic weather predictions which are sharp, calibrated and reliable both for central values and dispersion (see also Hagedorn et al., 2007; Hamill et al., 2007).

The paper is organized as follows. Section 2 illustrates the hierarchical model linking observed data and deterministic forecasts and points the way to obtaining the predictive distribution. Details of the model specification are put off until certain preliminary descriptive analyses of data have been completed at the beginning of Section 3: Section 3.4 completes the model. Section 4 illustrates the results and compares them with the performance of EPS models. Section 5 presents our conclusions.

2 Statistical modelling of EPS: a hierarchical approach

We propose the full Bayesian modelling of EPS output by means of a hierarchical model. For each day $t$ we have (possibly multivari-
ate) observed data $y_t$, together with $K$ distinct forecast ensemble members $X_{tk}$. Observed data and EPS outputs can be seen as the components of a statistical model where observed meteorological values are explained by EPS outputs considered as exogenous variables. The principal definitions are summarised in table 1.

We consider forecast scenarios for a particular day as random replications of a single data-generating process. In order to link the ensemble of replicated scenarios to a single observation, we introduce a latent process which randomly selects one scenario from the ensemble. In this way, we model the observed data conditionally on the selection process.

This situation is then translated into a hierarchical model, where the first level concerns observed values as a function of the selected deterministic forecast, while the second level governs the selection of this element.

2.1 Model construction

As we have already said, the EPS output on day $t$ consists of $K$ forecast scenarios:

$$X_t = \{X_{tk}; k = 1, \ldots, K\}$$

These scenarios are unlabelled and for each day there are $K$ new replications, each independent of the previous ones, characterised by a situation of exchangeability.

The first level of the model, i.e. the observation level, where the measurement error occurs, is written as:

$$p(y_t|Z_t, X_t, \theta) = f(X_t, Z_t; \theta)$$  \hspace{1cm} (2)

where the variable $Z_t$ selects the deterministic forecast for day $t$. This means that the distribution $f(X_t, Z_t; \theta)$ is not a function of all the weather scenarios, but only of the selected ensemble member. In (2) the observed data $y_t$ is a function of just one of the $K$ scenarios on day $t$. 

5
The scenario selection process is modelled by a latent stochastic process \( \{ Z_t \} \) which selects, from one day to the next, one scenario from among the \( K \) available scenarios, such that the observation on day \( t \) is considered as a random fluctuation around the selected weather scenario. As a result of the way in which weather scenarios are generated, there is no correlation in time. We propose to model the latent selection process \( \{ Z_t \} \) in the second level of the model, i.e. the process level, as i.i.d. with discrete uniform distribution:

\[
p(Z_t = k) = \frac{1}{K} \quad k = 1, 2, \ldots, K
\]  

(3)

such that, for any \( N \)-dimensional sequence of times \( t_1, t_2, \ldots, t_N \), the following relationship holds:

\[
p(Z_{t_1} = k_1, Z_{t_2} = k_2, \ldots, Z_{t_N} = k_N) = \prod_{j=1}^{N} p(Z_{t_j} = k_j)
\]

The selected weather scenario can be compactly indicated as a function of \( Z_t \) and \( X_t \):

\[
x_t = h(Z_t, X_t) = \sum_{k=1}^{K} 1_{\{Z_t = k\}} X_{tk}
\]  

(4)

Conditional on the selection of the weather scenario \( x_t \), data \( (y_t, X_t) \) is reduced to the (multivariate) pairs \( (y_t, x_t) \). Equation (2) can thus be rewritten with the help of (4) as:

\[
p(y_t | Z_t, X_t, \theta) = f(h(Z_t, X_t), \theta) = f(x_t; \theta)
\]  

(5)

In practice, the form of model (5) will depend on the available data set. The details of model specification are therefore deferred to the following Section. A standard choice would be Gaussian regression, with a data-driven specification of the regression function and of variance. Furthermore, if one wants to model a one-dimensional variable in multiple locations, \( y_t \) and \( x_t \) shall be \( S \)-variate vectors, and \( \theta \) shall contain, among other things, variance/covariance parameters
of the distribution of the $S$-variate vector of residuals. Statistical modelling according to (5) has calibration as a natural by-product, a required element which is absent from deterministic models.

We propose a Bayesian solution for model (5)-(3) whereby, after specifying a prior for $\theta$, one finds the posterior distribution $p(\theta|y^{\text{obs}})$, where $y^{\text{obs}}$ are the observed data from the training set.

In the following, conditioning on regressor $X_t$ will not be explicitly specified.

2.2 Predictive probability distribution

In the problem in question, we are particularly interested in the posterior predictive probability distribution of the target variable $y$ at a future point in time $T+1$, after observing $y^{\text{obs}} = \{ y_t; t = 1, \ldots, T \}$.

By standard Bayesian arguments, we can write:

$$p(y_{T+1}|y^{\text{obs}}) = \int_{\Theta} p(y_{T+1}|\theta)p(\theta|y^{\text{obs}})d\theta \quad (6)$$

A suitable expression for $p(y_{T+1}|\theta)$ can be obtained from (6) by conditioning w.r.t. $Z_{T+1}$ and applying the total probability law:

$$p(y_{T+1}|\theta) = \sum_{k=1}^{K} p(y_{T+1}|Z_{T+1} = k; \theta) \cdot p(Z_{T+1} = k)$$

$$= \frac{1}{K} \sum_{k=1}^{K} p(y_{T+1}|Z_{T+1} = k; \theta)$$

where the role of the first level of the model written according to (5) is explicitly expressed.

By substituting (7) in (6) we finally get:

$$p(y_{T+1}|y^{\text{obs}}) = \frac{1}{K} \sum_{k=1}^{K} \int_{\Theta} p(y_{T+1}|Z_{T+1} = k; \theta)p(\theta|y^{\text{obs}})d\theta \quad (8)$$

where the resulting distribution is a mixture of $K$ components with equal weights $1/K$. The integration in (8) may be hard to solve in analytical terms, but can be easily computed using numerical stochastic
methods. Note that in (8) the parameters vector $\theta$ is integrated out from the measurement error model (5), unlike in the BMA solution (1) which replaces $\theta$ with a point estimate $\hat{\theta}$ according to an empirical Bayesian approach.

3 Data-driven model completion

3.1 Forecast data: the COSMO-LEPS system

On the global scale, EPS is a technique commonly used to deal with forecast uncertainty, whereas small-scale use of EPS is still rare. Local meteorological centres generally compute point forecasts exclusively through high resolution Limited Area Models (LAM) applied to global forecasts given by international centres or use multimodel EPS for mimicking global EPS outputs (Krishnamurti et al., 1999; Kharin and Zwiers, 2002).

In 2001, the Emilia Romagna ARPA-SMR (Azienda Regionale Prevenzione e Ambiente - Servizio Meteorologico Regionale) developed LEPS (Local Scale Ensemble Prediction System) which, after an initial experimentation phase, lead to the implementation of COSMO-LEPS (Montani et al., 2001, 2003). The LEPS method employs global-scale ECMWF ensemble forecasts to obtain local-scale, high-resolution ensemble forecasts with relatively low computational costs (Marsigli et al., 2005).

LEPS uses ECMWF ensemble forecasts to obtain a reduced set of local forecasts. In other words, 10 representative members (RM) are selected from the set of 51 ECMWF forecasts by means of an unsupervised cluster analysis. Each RM gives starting and bound

<table>
<thead>
<tr>
<th>variable</th>
<th>description</th>
</tr>
</thead>
<tbody>
<tr>
<td>$y_t$</td>
<td>observed data on day $t$</td>
</tr>
<tr>
<td>$X_{tk}$</td>
<td>$k$-th weather scenario on day $t$, $k = 1, \ldots, K$</td>
</tr>
<tr>
<td>$x_t$</td>
<td>selected weather scenario on day $t$</td>
</tr>
</tbody>
</table>
conditions for the integration of a high-resolution LAM. So, a corresponding local ensemble member is obtained from each RM.

We consider temperature data at an altitude of 2m at 12:00, based on deterministic COSMO-LEPS forecasts with a forecast horizon of 24h.

3.2 Observed data

Observed data covers the period from 1/06/2005 to 30/11/2005, and comes from 12 meteorological stations belonging to the Italian Synoptic Network.

The positioning of the selected stations is reported in Fig. 1, while in table 2 geographic details are summarised. In table 3 some descriptive statistics of the observed data are synthesised. From geographic data as well as descriptive statistics, it can be seen how the selected stations are homogeneous w.r.t. orography, distance from sea and overall mean temperature. There are however some differences in the levels of variability, as indicated by the computed standard deviation (sd) and inter-quartile range (IQR) in Gela and Lampedusa, where they are noticeably smaller than in the other stations.
### Table 2: Details of the meteorological stations

<table>
<thead>
<tr>
<th>label</th>
<th>name</th>
<th>lon.</th>
<th>lat.</th>
<th>alt. (m)</th>
<th>missings</th>
</tr>
</thead>
<tbody>
<tr>
<td>Cvt</td>
<td>Civitavecchia</td>
<td>11.83</td>
<td>42.03</td>
<td>3</td>
<td>7</td>
</tr>
<tr>
<td>PDM</td>
<td>Pratica Di Mare</td>
<td>12.43</td>
<td>41.65</td>
<td>6</td>
<td>6</td>
</tr>
<tr>
<td>Grz</td>
<td>Grazzanise</td>
<td>14.07</td>
<td>41.06</td>
<td>9</td>
<td>6</td>
</tr>
<tr>
<td>Pnz</td>
<td>Ponza</td>
<td>12.95</td>
<td>40.92</td>
<td>184</td>
<td>7</td>
</tr>
<tr>
<td>Cpr</td>
<td>Capri</td>
<td>14.20</td>
<td>40.55</td>
<td>160</td>
<td>6</td>
</tr>
<tr>
<td>CpP</td>
<td>Capo Palinuro</td>
<td>15.28</td>
<td>40.03</td>
<td>184</td>
<td>5</td>
</tr>
<tr>
<td>MDG</td>
<td>Marina Di Ginosa</td>
<td>16.88</td>
<td>40.44</td>
<td>2</td>
<td>6</td>
</tr>
<tr>
<td>Ust</td>
<td>Ustica</td>
<td>13.18</td>
<td>38.71</td>
<td>250</td>
<td>6</td>
</tr>
<tr>
<td>TrB</td>
<td>Trapani Birgi</td>
<td>12.50</td>
<td>37.92</td>
<td>7</td>
<td>5</td>
</tr>
<tr>
<td>Gel</td>
<td>Gela</td>
<td>14.22</td>
<td>37.08</td>
<td>11</td>
<td>6</td>
</tr>
<tr>
<td>Lmp</td>
<td>Lampedusa</td>
<td>12.60</td>
<td>35.50</td>
<td>16</td>
<td>11</td>
</tr>
<tr>
<td>CpC</td>
<td>Capo Carbonara</td>
<td>9.517</td>
<td>39.10</td>
<td>116</td>
<td>11</td>
</tr>
</tbody>
</table>

### Table 3: Summary of observed temperature data

<table>
<thead>
<tr>
<th>station</th>
<th>mean</th>
<th>median</th>
<th>sd</th>
<th>IQR</th>
</tr>
</thead>
<tbody>
<tr>
<td>Cvt</td>
<td>23.12</td>
<td>24.00</td>
<td>4.662</td>
<td>6.00</td>
</tr>
<tr>
<td>PDM</td>
<td>23.18</td>
<td>24.60</td>
<td>5.139</td>
<td>6.40</td>
</tr>
<tr>
<td>Grz</td>
<td>24.44</td>
<td>26.00</td>
<td>5.913</td>
<td>7.60</td>
</tr>
<tr>
<td>Pnz</td>
<td>22.12</td>
<td>24.00</td>
<td>4.791</td>
<td>6.60</td>
</tr>
<tr>
<td>Cpr</td>
<td>23.77</td>
<td>25.00</td>
<td>5.655</td>
<td>8.20</td>
</tr>
<tr>
<td>CpP</td>
<td>24.27</td>
<td>26.00</td>
<td>5.608</td>
<td>7.65</td>
</tr>
<tr>
<td>MDG</td>
<td>24.69</td>
<td>26.00</td>
<td>5.649</td>
<td>7.80</td>
</tr>
<tr>
<td>Ust</td>
<td>24.37</td>
<td>24.60</td>
<td>4.71</td>
<td>6.20</td>
</tr>
<tr>
<td>TrB</td>
<td>25.74</td>
<td>26.10</td>
<td>4.677</td>
<td>5.60</td>
</tr>
<tr>
<td>Gel</td>
<td>24.65</td>
<td>25.00</td>
<td>3.62</td>
<td>3.90</td>
</tr>
<tr>
<td>Lmp</td>
<td>25.80</td>
<td>26.20</td>
<td>2.738</td>
<td>2.60</td>
</tr>
<tr>
<td>CpC</td>
<td>25.18</td>
<td>26.40</td>
<td>4.618</td>
<td>5.60</td>
</tr>
</tbody>
</table>
3.3 Explorative data analysis

Explorative evaluations are useful for suggesting the final details of model building. The differences between observed data and mean ensemble forecasts:

$$\bar{e}_{ts} = y_{ts} - \frac{1}{10} \sum_{k=1}^{10} X_{tsk}$$

have been considered and reported in Fig. 2 for each meteorological station. Such distributions are roughly symmetric, with comparable ranges. However, there is a positive bias of varying amplitude for all stations, except perhaps for Grazzanise. Further parameters will be introduced during model building in order to account for these differences between stations.

A more detailed view of the relationship between observed data and mean deterministic forecasts is given in Fig. 3. This plot would seem to indicate a linear relationship. On the basis of figure 2, a station-specific intercept should be added, while the slope can be considered approximately constant among stations.

To sum up, our explorative plots suggest:

- the need for a correction for a station-specific positive bias
Figure 3: Mean deterministic forecasts vs observed data, by station
a linear relationship, with nearly constant slope, between deterministic forecasts and observed data (Fig. 2 and 3)

- approximately constant variance (box-plot width in Fig. 2)

### 3.4 Statistical model completion

Now we are able to specify model (2) by assuming that \( x_t \) and \( y_t \) are \( S \)-variate vectors:

\[
x_t = (x_{t1}, x_{t2}, \ldots, x_{ts}, \ldots, x_{tS}) \]
\[
y_t = (y_{t1}, y_{t2}, \ldots, y_{ts}, \ldots, y_{tS})
\]

and complete the specification of the regression function guided by the exploratory analyses in Section 3.3.

We assume for \( y_{ts} \), conditional on \( x_{ts} \) and \( \theta \), a Gaussian distribution with expected value constituted by a linear function of the deterministic forecast \( x_{ts} \):

\[
y_{ts} | x_{ts}, \theta \sim N(\alpha_s + \beta x_{ts}, \sigma^2_y) \quad (9)
\]

where \( \beta \) is a common slope, \( \alpha_s \) are station-specific intercepts such that \( \alpha = (\alpha_1, \ldots, \alpha_S) \), and \( \sigma^2_y \) is a common measurement error variance. Here \( \theta_1 = (\alpha, \beta, \sigma^2_y) \), while the \( y_{ts} \) scalars are modelled as independent, both in time and space, conditionally on model parameters. In order to complete the model we give a hierarchical structure to the station-specific intercepts:

\[
\alpha_s | \alpha_0, \sigma^2_\alpha \sim N(\alpha_0, \sigma^2_\alpha) \quad (10)
\]

and assign vague priors to \( \beta, \alpha_0, \sigma^2_y \) and \( \sigma^2_\alpha \).

The model parameters consist of the vector \( \theta = (\theta_1, \theta_2) \), composed of 12 intercepts, 1 common slope, 1 common variance of the measurement error and 2 hyperparameters \( \theta_2 = (\alpha_0, \sigma^2_\alpha) \) for the distribution of intercepts.
4 Results

In order to enable a comparison with COSMO-LEPS ensembles, models have been fitted on a moving window of 30 days, and out-of-sample forecasts simulated for 1 day ahead. This procedure estimates 153 models, corresponding to about 5 months.

For each day, model estimation gives:

- the full joint posterior probability distribution of the parameter vector \( p(\theta|y^{\text{obs}}) \), from which we can obtain all lower-order marginals
- the predictive distribution function \( p(y_{T+1}|y^{\text{obs}}) \) computed in (6)

Since we estimate the model for 153 different training sets, the previous estimates are replicated for each target day. This procedure is in keeping with the current work of meteorologists, who look for daily evaluations of forecasts. Unless otherwise stated, for exploratory purposes we have focused on inferences made for the 1st July 2005, which are thus conditional on the 1 – 30 June 2005.

4.1 The posterior distribution of parameters

One of the results of model estimation is the multivariate posterior probability distribution \( p(\theta|y^{\text{obs}}) \). Among parameters, 12 are the station-specific intercepts \( \alpha_i \), and we do not discuss them since they are less interesting for the problem under study. In Fig. 4 the marginal and bivariate densities (as isodensity contours in the out-of-diagonal elements) for the remaining 4 parameters are reported.

Bivariate distributions show no pathological posterior dependency between parameters. In fact, pairwise independence holds for almost all parameters, according to the mild shape of the contours, the only exception being the \( (\alpha_0, \sigma_\alpha) \) distribution, which shows a form of scale-location dependency.

Parameters \( \beta, \sigma_y \) and \( \alpha_0 \) display (see the diagonal) a symmetric distribution: in the case of \( \beta \) and \( \alpha_0 \) this was to be expected, since
Figure 4: Univariate and bivariate parameter densities. Univariate density on the diagonal, bivariate contours out of diagonal.
they are location parameters; in the case of $\sigma_y$ such a shape is likely
due to the large amount of independent data that contributed to the
estimation ($12 \times 30$ observed-deterministically forecasted data pairs).
The distribution of $\sigma_\alpha$ is skewed, which is characteristic of dispersion
parameters. As regards posterior precision, both $\beta$ and $\sigma_y$ display
a very sharp marginal distribution, with $\beta$ roughly ranging from 0.8
to 1.0 and $\sigma_y$ from 1.3 to 1.7. Note that the distribution of slope
$\beta$ is positioned well below 1. In a certain sense, it is significantly
smaller than 1, thus confirming the need for the contraction of the
deterministic forecasts in order to obtain improved calibration. The
marginal distributions of $\alpha_0$ and $\sigma_\alpha$ are much more dispersed than
the others, but this is to be expected given that they are second-level
parameters.

Table 4 shows some summaries of the marginal posterior distribu-
tions. Here the above-mentioned remarks about Fig. 4 are confirmed
by the estimated standard deviations (sd) and inter-quartile ranges (IQR).

### 4.2 Predictive distributions

Fig. 5 plots the predictive probability densities (PD) for the monitor-
ing stations on the 1st July 2005. The observed value (solid vertical
line) and the range of the deterministic ensemble forecast (dotted
vertical lines) are shown in each plot. The PD is estimated on data
up until the 30th June. In this way, a fair comparison can be made
with the COSMO-LEPS ranges, since we are comparing the statisti-
cal forecast with data which did not contribute to model estimation.
Table 5: Observed and mean predicted values for the 1st July 2005

<table>
<thead>
<tr>
<th>station</th>
<th>observed</th>
<th>EPS mean</th>
<th>PDF mean</th>
</tr>
</thead>
<tbody>
<tr>
<td>Cvt</td>
<td>27.6</td>
<td>26.5</td>
<td>28.5</td>
</tr>
<tr>
<td>PDM</td>
<td>28.2</td>
<td>26.3</td>
<td>28.3</td>
</tr>
<tr>
<td>Grz</td>
<td>28.8</td>
<td>28.6</td>
<td>28.9</td>
</tr>
<tr>
<td>Pnz</td>
<td>25.8</td>
<td>24.6</td>
<td>25.9</td>
</tr>
<tr>
<td>Cpr</td>
<td>32.0</td>
<td>25.3</td>
<td>28.3</td>
</tr>
<tr>
<td>CnP</td>
<td>29.0</td>
<td>26.6</td>
<td>28.5</td>
</tr>
<tr>
<td>MDG</td>
<td>32.0</td>
<td>26.2</td>
<td>28.2</td>
</tr>
<tr>
<td>Ust</td>
<td>28.4</td>
<td>25.0</td>
<td>27.7</td>
</tr>
<tr>
<td>TrB</td>
<td>29.0</td>
<td>24.8</td>
<td>27.3</td>
</tr>
<tr>
<td>Gel</td>
<td>27.6</td>
<td>27.2</td>
<td>28.2</td>
</tr>
<tr>
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<td>28.9</td>
<td>24.9</td>
<td>26.7</td>
</tr>
<tr>
<td>CpC</td>
<td>29.4</td>
<td>25.3</td>
<td>28.3</td>
</tr>
</tbody>
</table>

Table 5 compares observed data and mean predicted values.

The observed values at the various stations may or may not fall within the COSMO-LEPS range, hence the significant contribution made by the statistical model to deal with the weather forecasting problem. In fact, when the observed data falls within the EPS range (Grz and Gel), the PD mean also lies within that range. When the observed data falls outside the said range, the PD tends to indicate those values that fall between the deterministic COSMO-LEPS range and the actually observed data as the most probable (Cpr, CnP, MDG, Ust, TrB, Lmp and CpC). In some cases (Cvt, PDM and Pnz), the PD emphasises the differences noticed between the COSMO-LEPS range and the observed value, but with more weight given to values closer to the observed value. In all cases, the overall tendency sees the PD assume the highest values that are closest to the observed value (but that have not been used for model estimation).

Note that the forecast calibration is not an explicit target of model estimation, but is the most important by-product of the fitting process, and one that does not require any form of human tuning.

In Fig. 6 we report the set of daily Ustica PD computed for July-August. The darker area indicates the central 50% Highest Density
Figure 5: Observed data, the COSMO-LEPS range and model forecast for the 1st July 2005
Figure 6: July-August HDI (coloured area), COSMO-LEPS range (dotted lines) and observed data (continuous line) for Ustica

Interval (HDI), while the lighter area represents the 90% HDI. The dotted lines indicate the EPS range, while the continuous line represents the observed data. This plot once again reveals the shift of the model-based PD towards the observed data for the 2 months in question, as well as the good performance of the PD in relation to both the central value and dispersion.

4.2.1 Verification Rank Histograms

In ensemble forecasting practice, certain standard diagnostic measures are employed to evaluate the performances of EPS. These measures can be applied to the output of our statistical model, thus enabling a direct comparison to be made with the COSMO-LEPS result.

A common diagnostic device is the Verification Rank Histogram (VRH; see Talagrand et al., 1997). The VRH is the histogram of frequencies of the rank of the observed data within the forecast ensemble. A good EPS should have a uniform distribution, meaning exchangeability between deterministic predictions and observed data (Buizza, 1997; Buizza et al., 2005; Gneiting et al., 2007). Other shapes may indicate under/overdispersion and bias. In particular, a marked U-shaped distribution means that observed values tend to
be systematically outside (above or below) the extremes of the EPS range. A concave shape indicates that the forecasting range is larger than what is actually needed.

Fig. 7 shows the VRH for the full dataset, both for the statistical model and for the deterministic EPS. It shows a clear lack of calibration in the case of deterministic EPS, while the results coming from the statistical models present a rather uniform distribution, indicating a good mix of model-based out-of-sample forecasts and actually observed data.

If we compute station-specific VRHs, we do not obtain as good results as for the overall histogram. Less data is available for each VRH, and the overall histogram is an average of the local evaluations. In each station-specific VRH, the shape is however far from the U-shape displayed by the global COSMO-LEPS VRH (see Fig. 8).

4.3 MCMC diagnostics

The model built in Sections 2 and 3 cannot be estimated analytically, but requires the use of numerical methods. MonteCarlo Markov Chains (MCMC) are a natural choice for the estimation of hierarchical models. MCMC can be problematic with respect to chain convergence and mixing (Robert and Casella, 1999). However, with regard to the problem in question, thanks to the adoption of conjugate priors
Figure 8: Station-specific statistical model VRH
for all parameters, which give standard full conditional distributions, as in Gelman 2004, genuine Gibbs sampling can be used. The Gibbs sampler performed well in this particular case.

So, for each daily model estimation, 5000 MCMC iterations were kept after discarding the first 1000. From each estimated model, 1000 scenarios were simulated from the predictive distribution. Chains revealed fast convergence and good mixing properties, with no need for ad-hoc fine tuning.

For the sake of example, we show some MCMC diagnostics for the model fitted in the period 1-30 June 2005 and used on the 1st July 2005. Fig. 9 shows MCMC traces, plotting sampled values vs. iteration number. These plots show a rapid convergence towards the target stationary distribution, with no real need to cut out any starting transient. Chains mixing is assessed here by means of a linear autocorrelation index. Fig. 10 plots sample autocorrelations at 36 different lags for the 4 main model parameters. Autocorrelations are almost null at all lags for $\alpha_0$, $\sigma_y$ and $\sigma_\alpha$, and negligible at lags $>3$ for $\beta$. This guarantees that the sampling process is highly efficient, so that there is no need for many replicates in order to reliably estimate parameters distribution.

5 Conclusions

EPS try to account for uncertainty due to partial knowledge of starting conditions, but without explaining residual variability. This study presents a genuinely Bayesian framework designed to deal with EPS constituted by unidentifiable members. An explicit link between EPS output and observed data is posited by introducing a latent selection process. We illustrate one application to small-scale temperature data for 12 meteorological stations. The model’s output is the full multivariate posterior probability distribution of the set of parameters characterizing the model, from which the following marginal, as well as bivariate, syntheses can be computed: central values, standard deviations, cross-correlations, credibility intervals, probability of falling in fixed ranges, etc. Our probabilistic model also allows us
Figure 9: MCMC traces

Figure 10: MCMC autocorrelations
to compute reliable, informative predictive distributions.

Since a Bayesian model produces a wealth of information, its results can be interpreted in a variety of different ways. One way of decoding such results is the combined use of synthetic values of the posterior with the predictive distributions. The description of our model’s results in fact makes use of several such syntheses, combining them in a flexible way, thus showing how a statistical model can help understand and improve deterministic models. In more specific terms, the predictive distribution obtained by the statistical model tends to be nearer to the observed values than does the COSMO-LEPS output.

Unfortunately, statistical modelling does not reduce the variability of results. Perhaps a more sophisticated form of modelling could represent a step in this direction. Indeed, statistical models tend to be rather simplistic, and their added value is constituted by their ability to bring together the various different components of the problem in hand in a suitable, consistent manner. Among such components, account is taken of the sources of variability in the right place. Finally, one natural by-product of statistical models is calibration, which is often required.

References


